



Progress on AST3-3 IR Camera Design

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AST3-3 IR Camera and KISS

- AST3-NIR is version of AST3 dedicated to conduct the KISS infrared sky survey at Kdark.
- Technical concept developed during early 2014.
- China responsible for telescope hardware and control, logistics, deployment.
- Australia responsible for instrument hardware and control, and power generation system.
- AST3-3 telescope already funded and currently is at manufacturing. IR camera funded by Australian government LIEF grant 2014(PI:Jeremy Mould) .
- Project Kick-off meeting happened during 2015 International Collaboration Meeting on Antarctic Survey Telescopes (AST3), March, Hong Kong 2015.

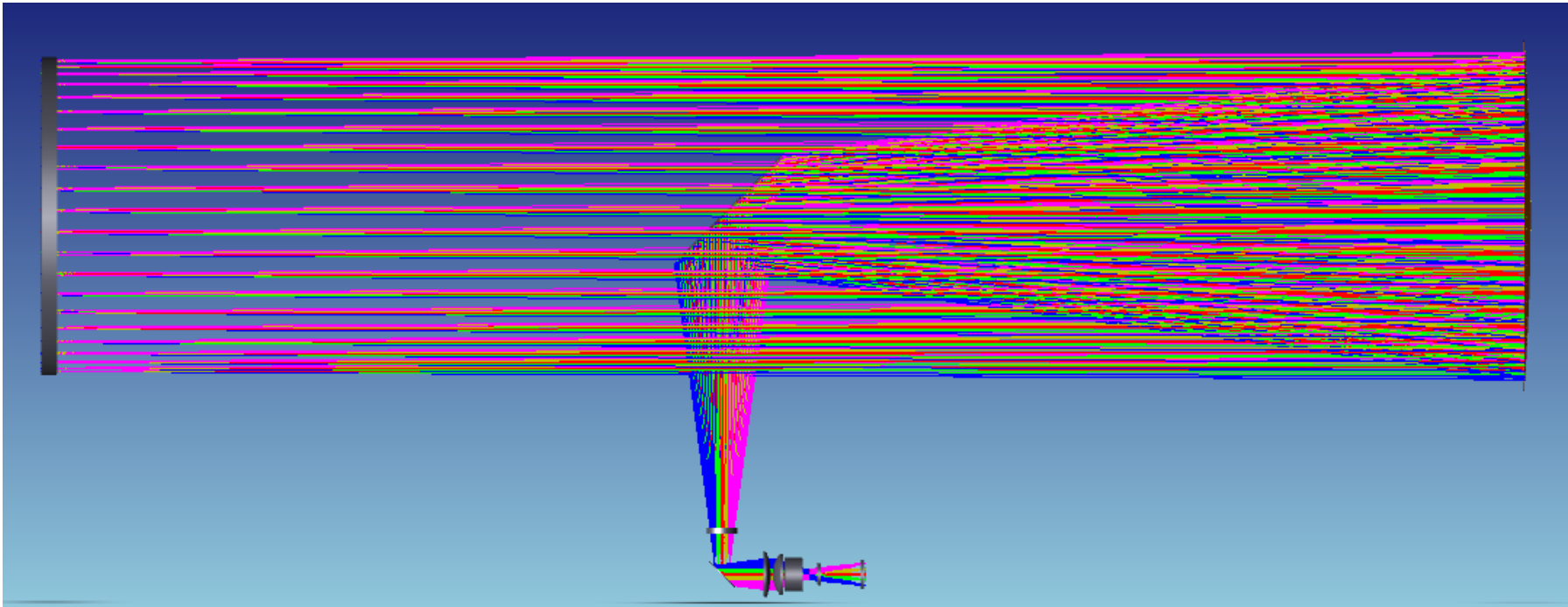


Key parameters (thanks to Jeremy, Jon and Xiangyan)

Parameter	Value
λ ($\Delta\lambda$)	2.36 (0.18) μm (K_{dark})
Diameter	68cm
Image Quality	1.9" (1.1 x diffraction limit)
Array	2048 x 2048, 18 μm pixels H2RG Teledyne preferred
Sampling	1"
Field of View	30' x 30'
<i>Achieving:</i>	
Background limited integration time	25 secs
1 σ 25 seconds	18.0 mags.
10 σ 1 hour	18.5 mags.
Saturation limit (in 25 sec)	$K_{\text{dark}} = 11.1$ mags.
With Background Sky [South Pole]	$K_{\text{dark}} = 17.0$ mags/arc ² =100 μJy /arc ²

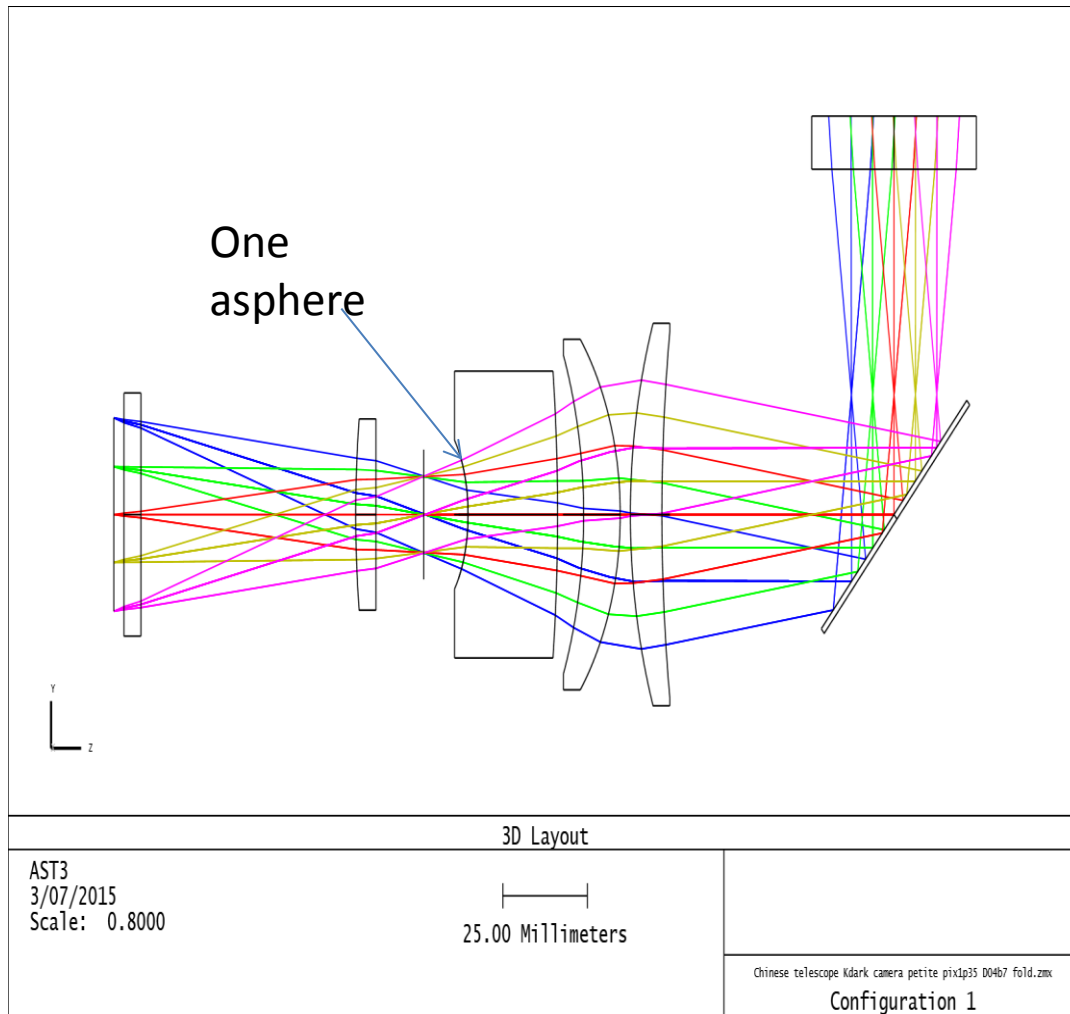


AST3-3 IR Camera Preliminary Design





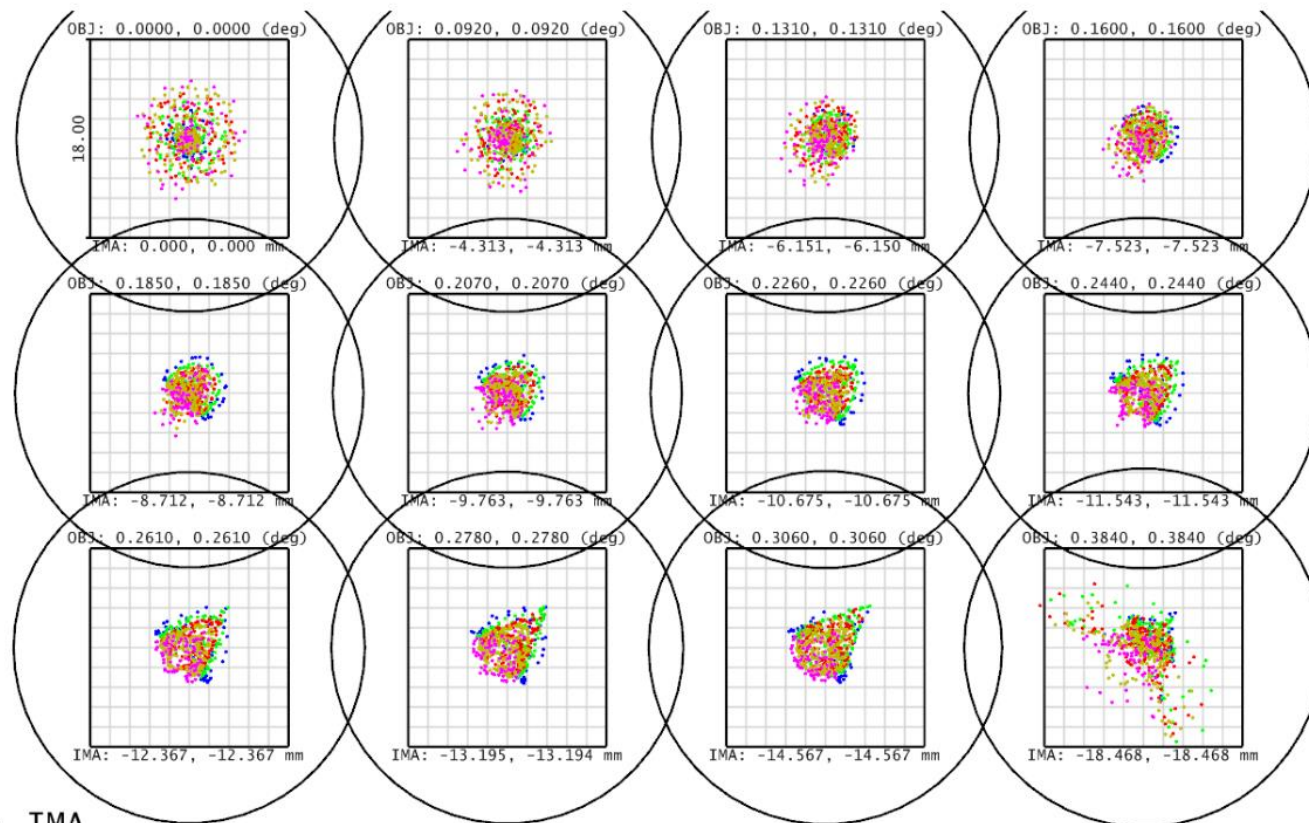
Optical design of IR camera



- Design has one asphere.
- Pixel size:
 - ✓ 18 μm .
 - ✓ 1.35" based on approximate calculations of size of total PSF including diffraction, as-designed aberrations, tolerances and seeing.
- Field:
 - ✓ 2048 x 2048 pixels.
 - ✓ 2765" x 2765".



PSF



Surface: IMA

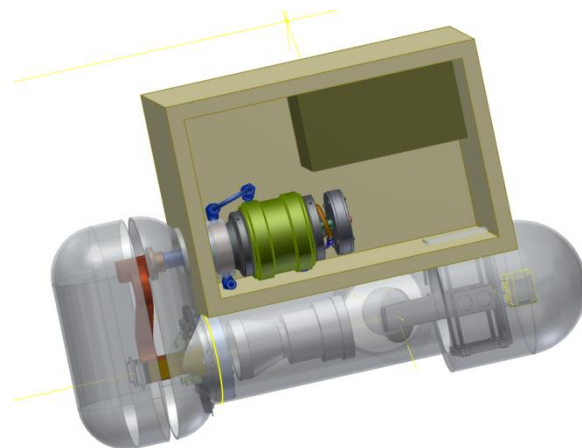
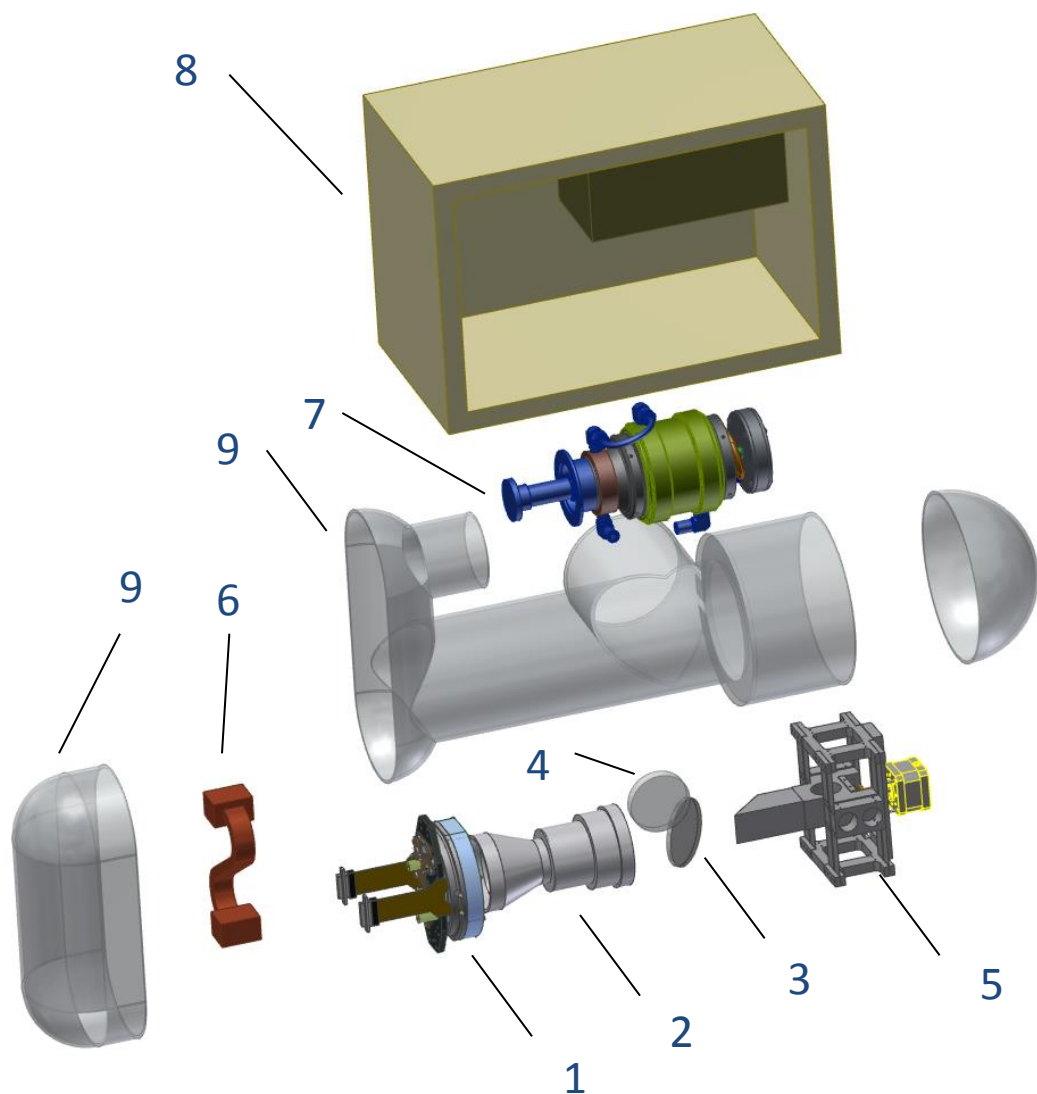
Spot Diagram

AST3												
3/07/2015												
Units are μm .												
Field :	1	2	3	4	5	6	7	8	9	10	11	12
RMS radius :	2.256	1.876	1.691	1.608	1.666	1.772	1.900	1.993	2.114	2.186	2.228	3.098
GEO radius :	5.575	4.937	4.632	4.044	4.060	3.915	4.098	4.199	5.052	5.051	5.549	10.097
Box width :	18	Reference : Centroid										

Chinese telescope Kdark camera petite pixIp35 D04b7 PSF.zmx
Configuration 1 of 1



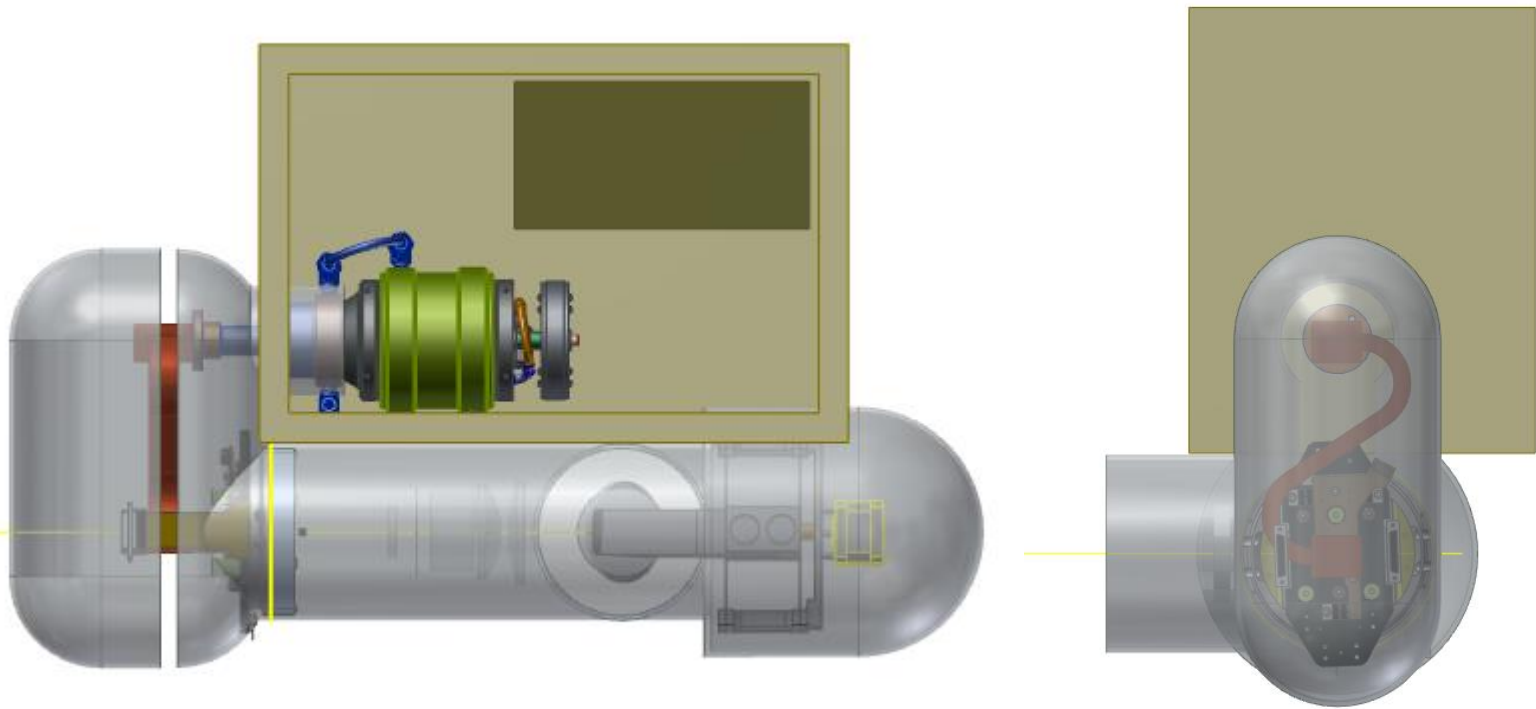
IR camera Concept



- 1 – Detector assembly
- 2 – Optical Camera
- 3 – mirror
- 4 - Vacuum window
- 5 – Focusing Mechanism
- 6 – Cold Strap
- 7 – Cryocooler (Sunpower Stirling)
- 8 – Electronics Thermal Cabinet
- 9 – Vacuum vessel



AST3-3 IR Camera Layout

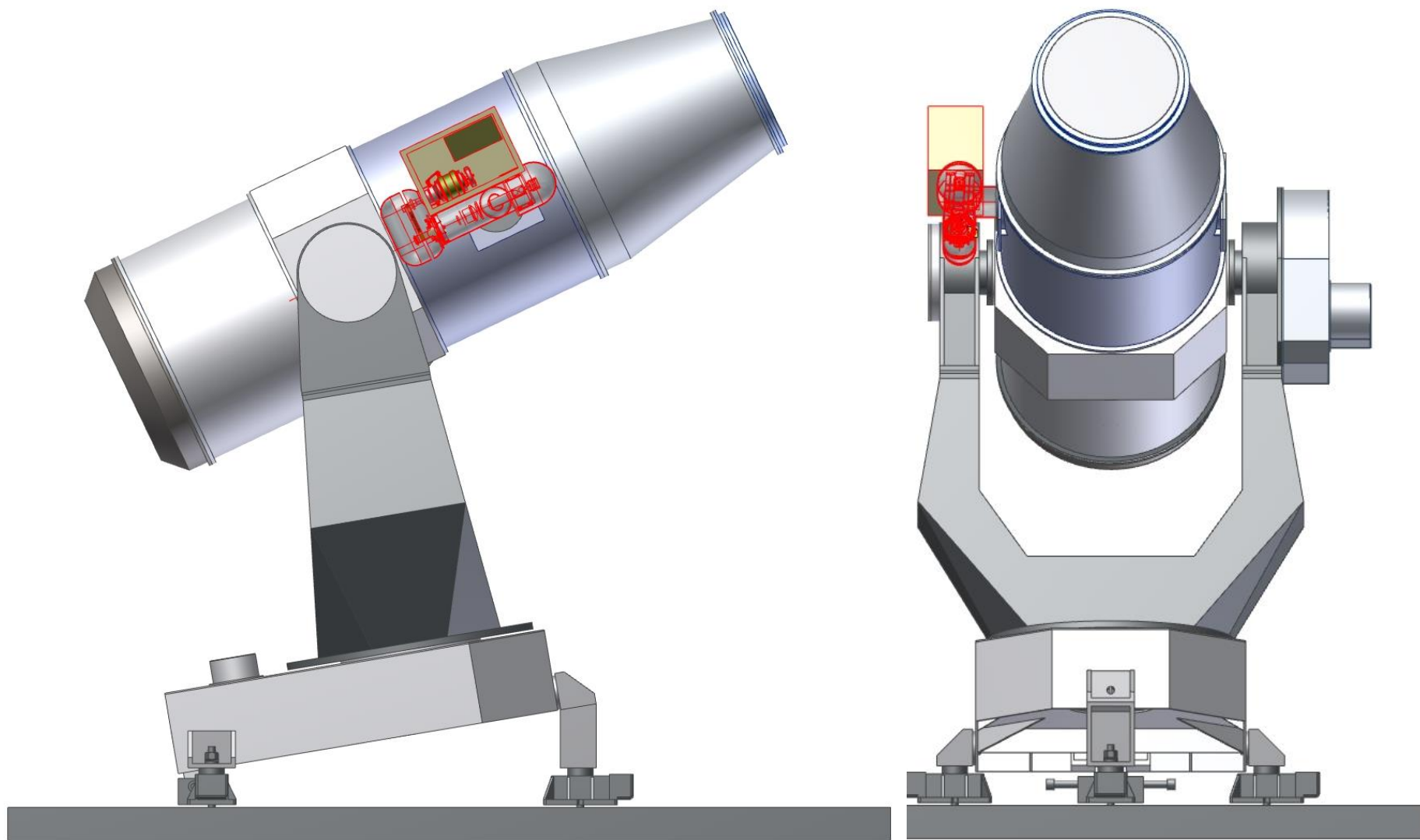


Overall Dimensions, approx: 700 x 500 x 300

Mass approx 30 kg



IR Camera on AST-3 telescope





Stray light and thermal analysis

- Scattering light analysis
- Thermal self emission

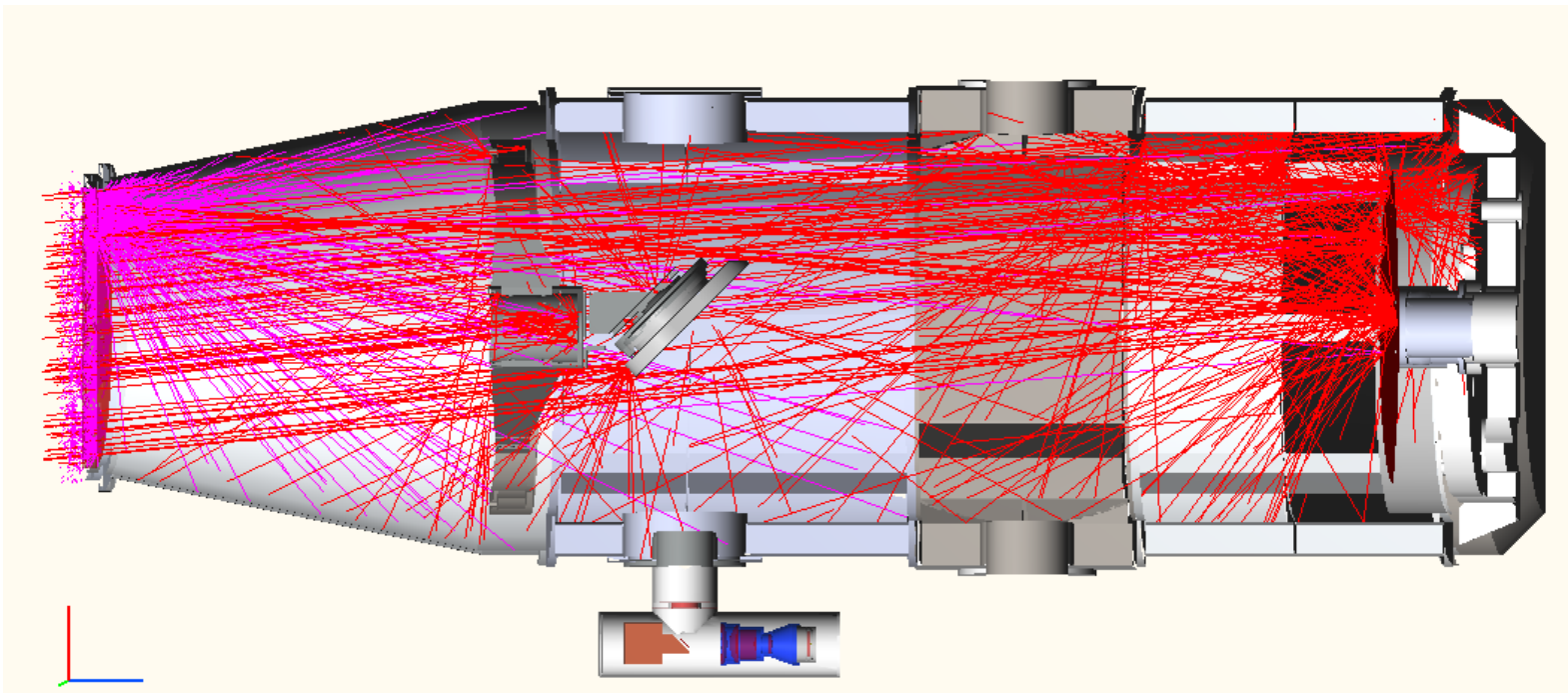
Surface property parameters(thanks to NIOAT)

material	absorptance	Specular Reflectance	Transmittance	BRDF A	BRDF B	BRDF g	Emissivity
mirror	0.03	0.956	0	0.00096	0.01	2	0.03
lens	0.001	0.015	0.984	0.00001	0.015	2	0.001
Tube/baffle	0.95	0.00001	0	0.0318	1	0	0.95
ITO	0.13	0.00187	0.85	0.0001	0.015	2	0.13

Point Source Transmittance

PST: Standard for stray light rejection evaluation.

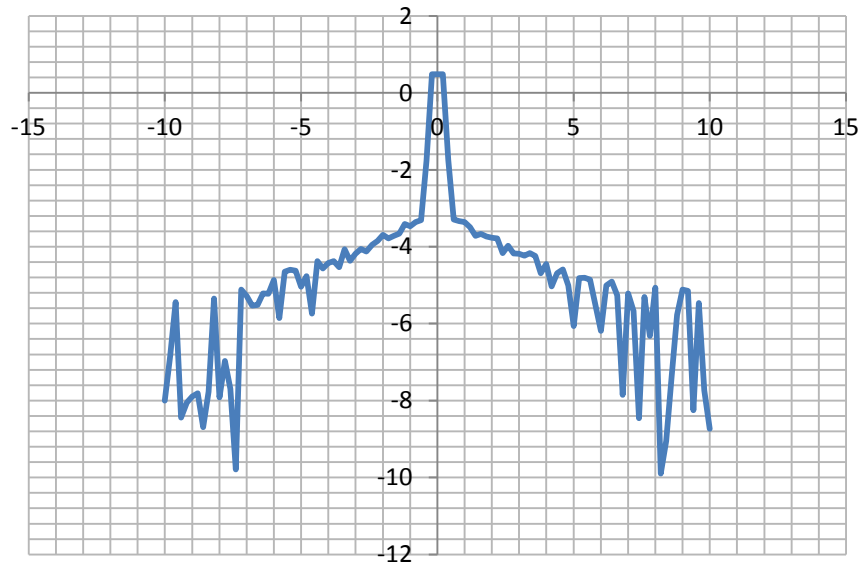
$PST(\theta) = \text{focal plane irradiance} / \text{entrance aperture irradiance}(\theta)$



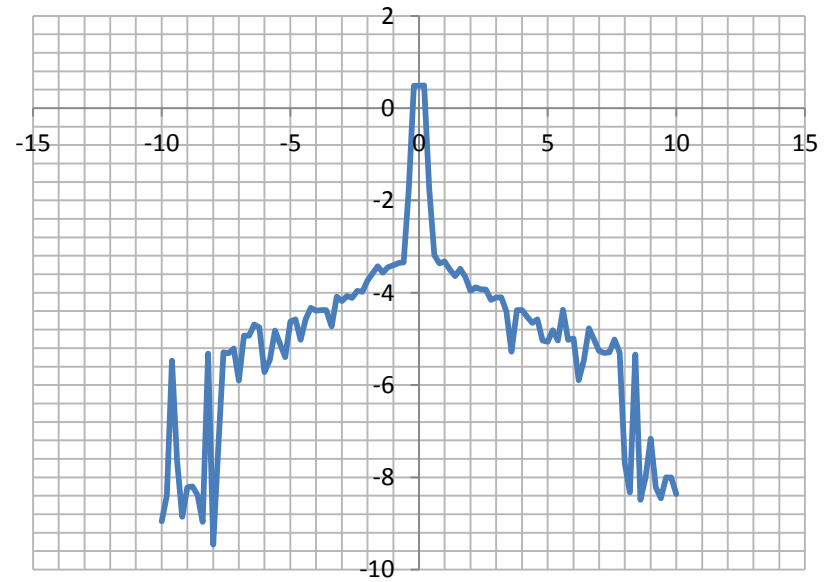


PST for IR camera

PST IR Camera_X direction

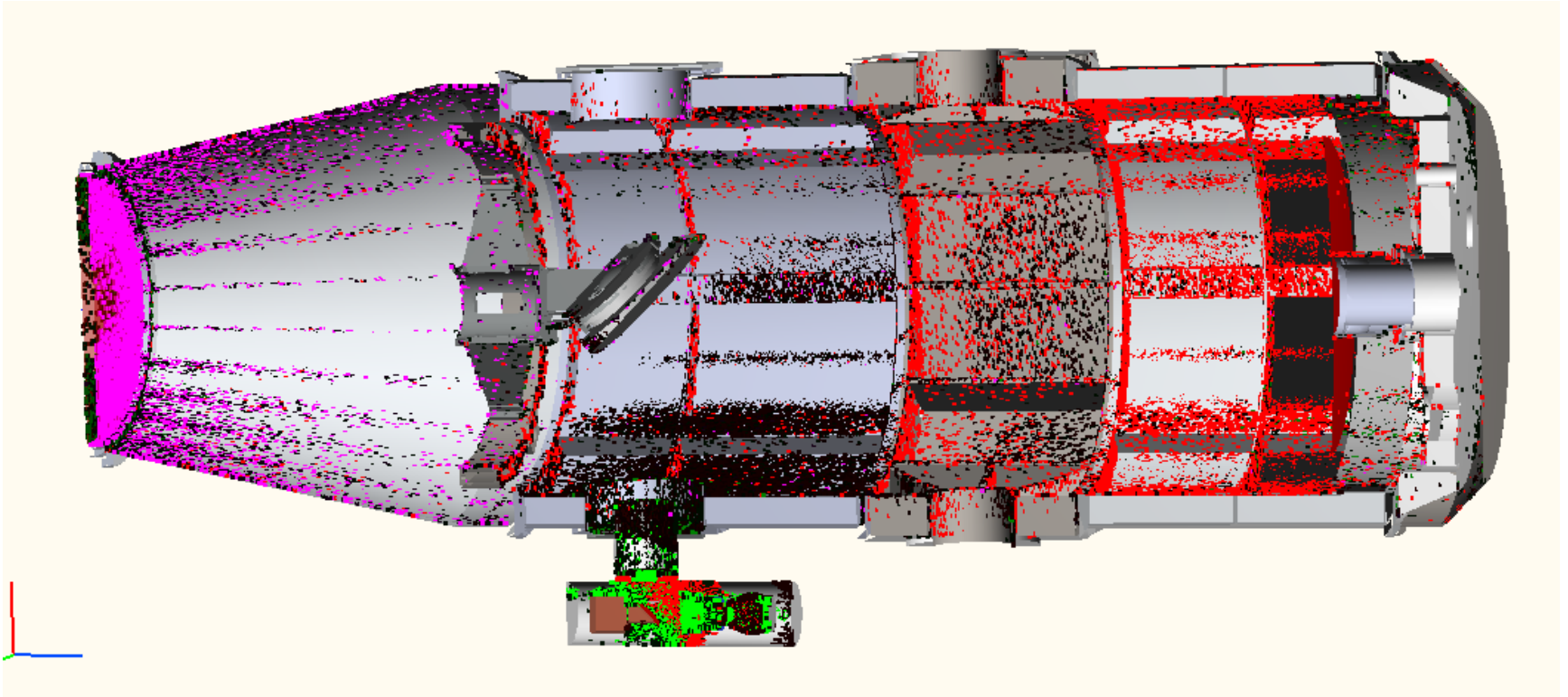


PST IR Camera_Y direction



Thermal self emission

Flooded Light from Detector Surface(Solid angle)

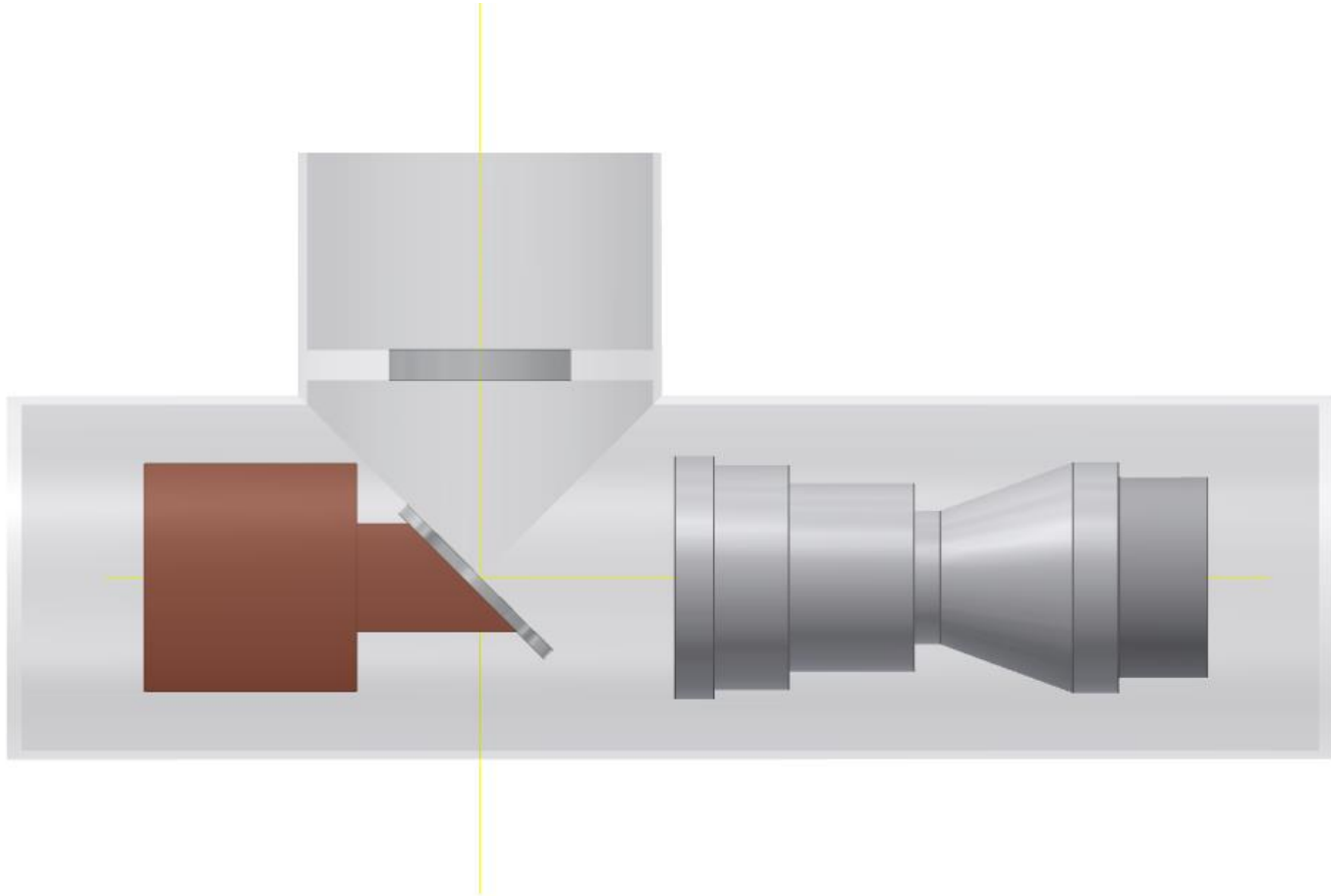


Three Temperature Zones:

- ✓ Zone1: Ambient temperature.
- ✓ Zone2: IR camera relay lens body.
- ✓ Zone3: Filter and FPA.



Model for stay light analysis

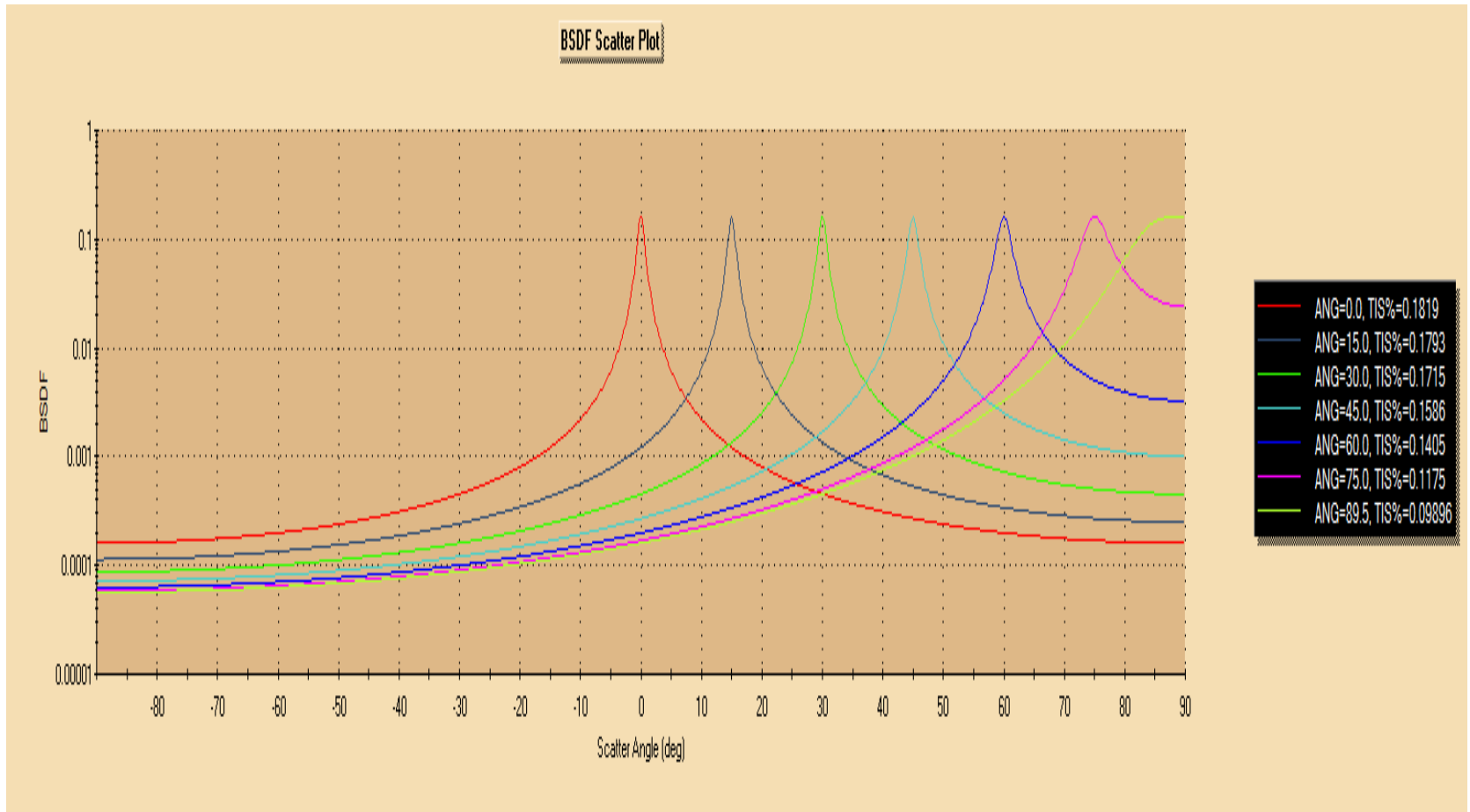




Thermal emission with Black paint

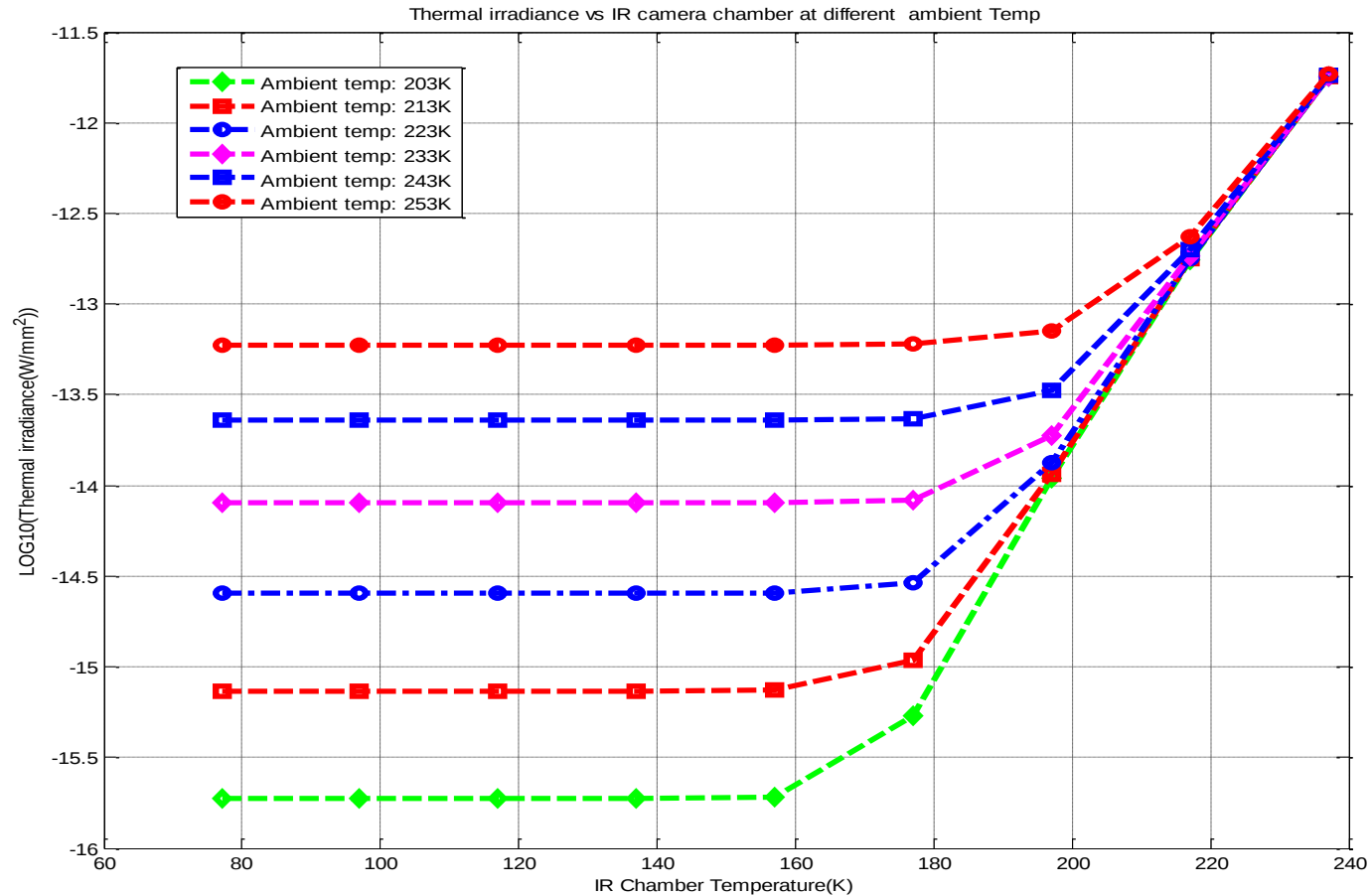
- Thermal irradiance with IR camera chamber temperature.

Black paint



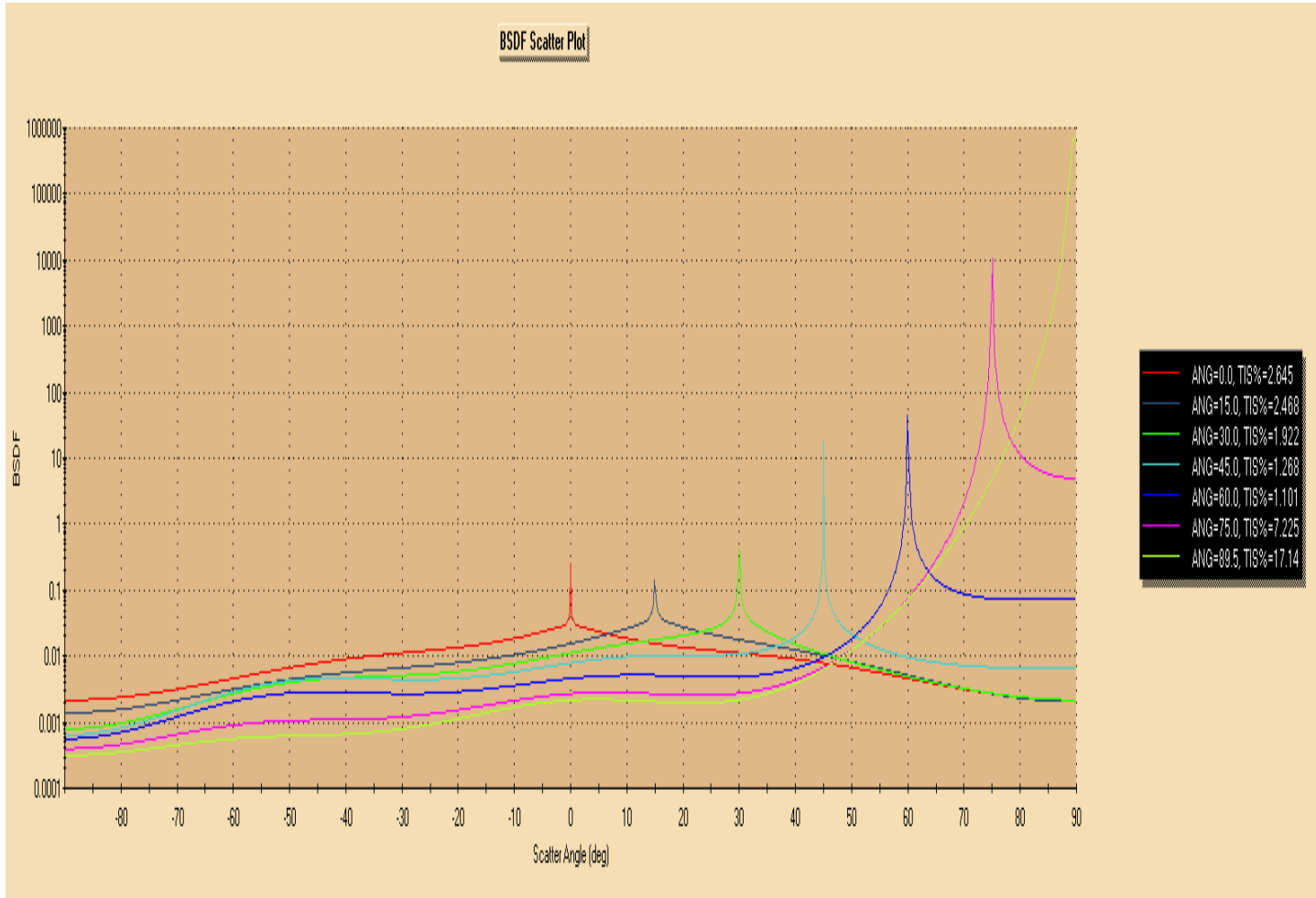


TSE vs IR Camera chamber temperature with black paint



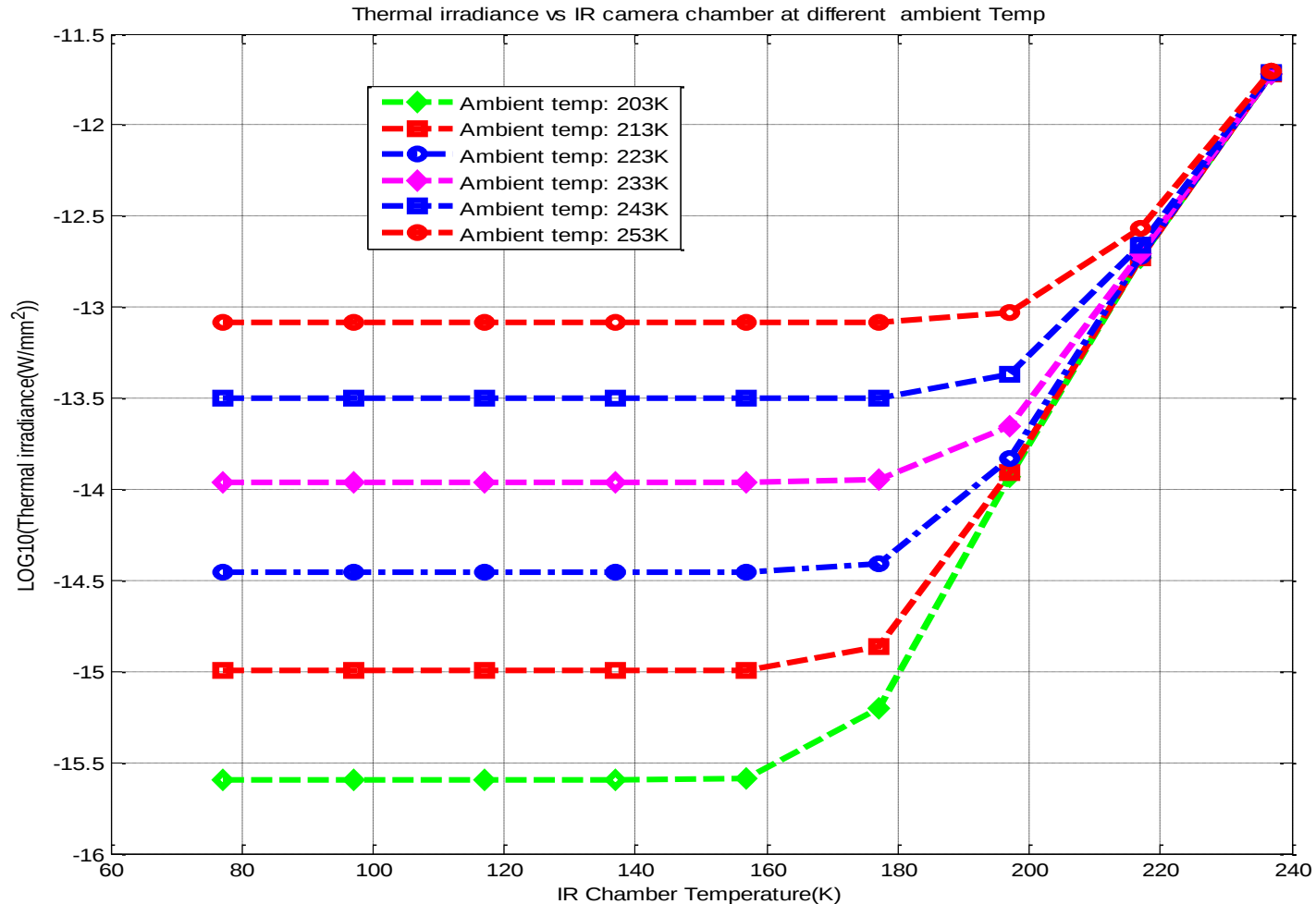


Acktar Vacuum black





Thermal irradiance when the IR camera chamber painted with Acktar vacuum black paint





AST3-3 AAO IR camera performance(no thermal)

AST3-3 IR Camera Exposure Time Calculator (V.0.1)

Target flux distribution

Object Magnitude: Observation Wavelength(um):

Target Geometry

☒ Point source ☐ Extended source:

CCD parameters

Pixel size: Dark Current: Read noise: Quantum Efficiency: Full Well Depth:

Sky conditions

Seeing: Sky Background: Mag: Jy

Instrument Setup

Filter Setup: Start Wavelength: Stop Wavelength: Ambient Temperature: Zenith Distance:

Signal to Noise Ratio(Defined by User)

Exposure Time(s): *Within Full Well Depth*

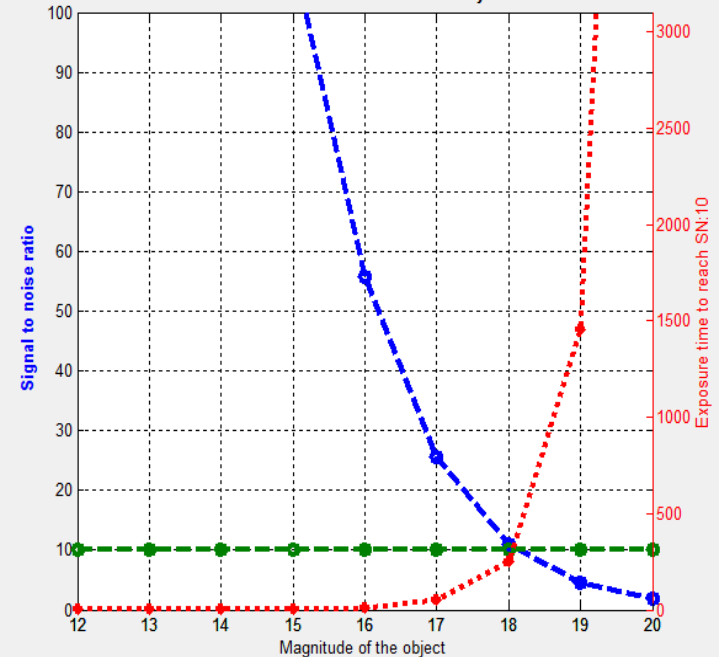
Maximum Exposure Time(s)

Results

Photons from Object: Photons from Sky Background: Photons from Telescope:

Exposure Time(s): SN_GivenTime:

SN vs magnitude of star, Observation time:300s.
Ambient temperature:0K. Zenith:0 degree.
AST3-3 IR Camera Preliminary

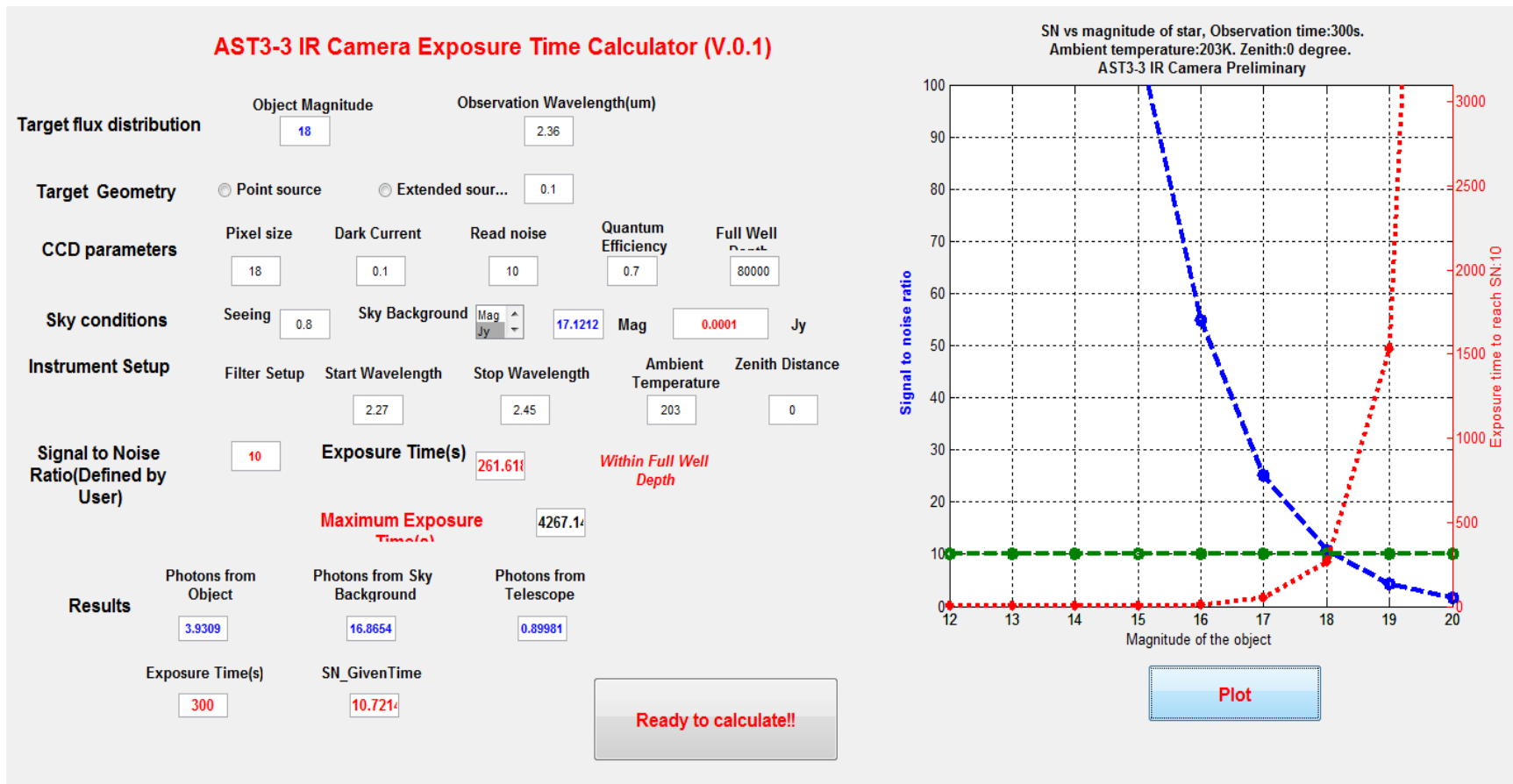


Plot

Ready to calculate!!



IR camera performance(with thermal: AmbientTemp: 203K)



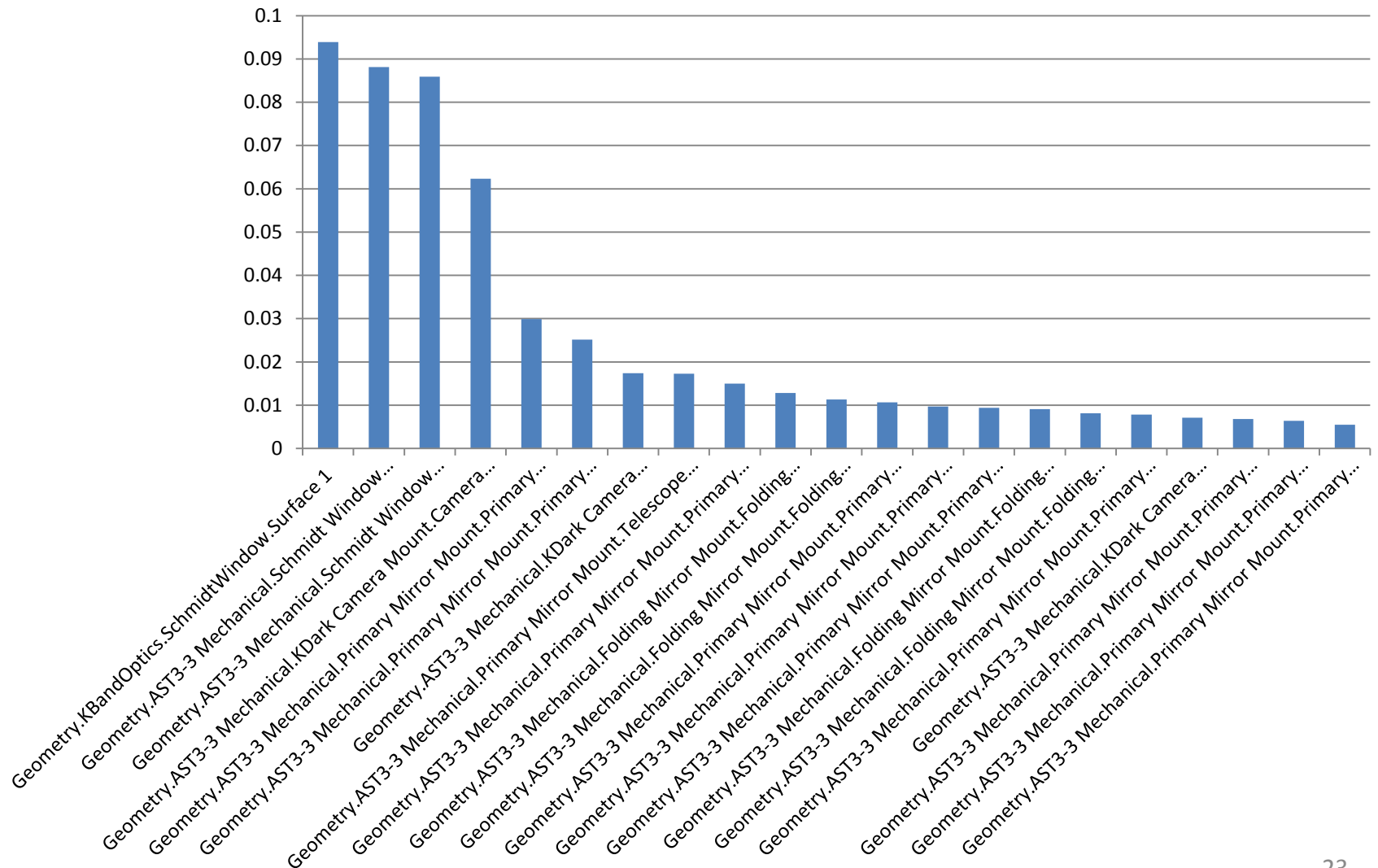


Find major parts contributed to thermal irradiance

- All IR camera chamber cools to $<160\text{K}$,
 - Schmidt window, major contributor
 - IR camera connector flange
- When IR camera chamber doesn't cool
 - IR Camera relay optics and all baffles dominate.

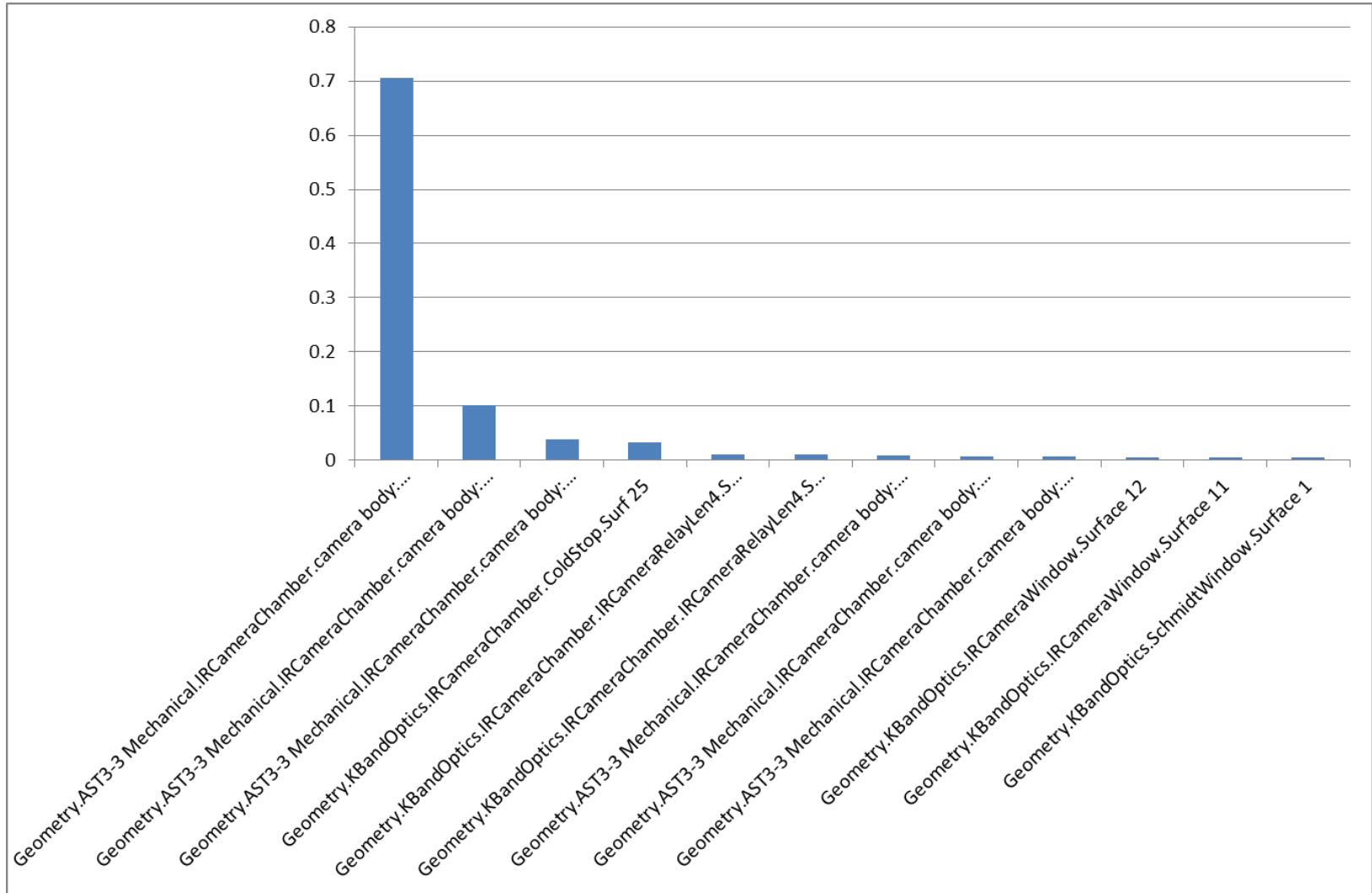


Major thermal contribution from Telescope and IR camera(all optics cool down to 77K(54%))





Major thermal contribution from Telescope and IR camera (Chamber temp to 197K(93.4%))





Plans

- Schmidt window thermal control. When telescope operate in IR, is it possible to defrost using other methods.



Key performance achieved with current design

- FOV: 0.584 degree^2
- 1σ 25 seconds: 18.5 mags at 235K
- 10σ 1 hour: 18.7 mags at 235K
- Kdark sky background: $100\mu\text{Jy}/\text{sec}^2$



More science requirements?

- Can we lock in our current design?

